Introduction to the sky
On a clear, moonless night, far from city lights, the night sky is magnificent. Roughly 2000 stars are visible to the unaided eye. If you know where to look, you can see Mercury, Venus, Mars, Jupiter, Saturn, and even Uranus. Occasionally, a bright comet is visible. On certain nights of the year there are many meteors (shooting stars) to be seen. Sometimes man-made debris falls back to Earth and burns up in the atmosphere.

Space Shuttle fuel tank reentry, April, 1984. Lava from Kilauea volcano illuminates clouds on the horizon.
We all know that the Sun rises in the east and sets in the west. The ancient Greeks attributed this to Apollo driving his chariot across the sky.

The Sun *appears* to complete a whole circle (360 degrees) around the Earth every 24 hours. Thus, its apparent motion is 15 degrees per hour from east to west. This apparent motion is due to the rotation of the Earth.

It turns out that the Sun can rise in the east-northeast, due east, or east-southeast depending on the day of the year.
Your *latitude* is equal to the number of degrees that the North Celestial Pole (near the star Polaris) is above the northern horizon.

If you’re close to the Earth’s equator, but still in the northern hemisphere, Polaris will be low in the sky, near the horizon.

If you’re above the Arctic Circle, Polaris will be high in the sky at all times.
Daily path of the Sun in the sky at latitude 23.5 deg N.

The Sun rises due east and sets due west only on the first day of spring and the first day of autumn. Note that the Sun at noontime on June 21st is 47 degrees higher in the sky than on December 21st.
The first day of spring occurs about March 20th or March 21st. (That's in the northern hemisphere.)

The first day of autumn occurs about September 22nd. (This would be the first day of spring in the southern hemisphere.)

The key thing is that the equinoxes occur when the declination of the Sun is zero degrees (i.e. located on the celestial equator).
Daily path of the Sun in the sky at the Earth's equator.

Note that the Sun rises due east and sets due west only on the first day of spring and the first day of autumn (about March 21st and September 22nd).
Daily path of the Sun in the sky at latitude 66.5 deg N. On June 21\textsuperscript{st} the Sun is 47 degrees higher in the sky at noontime than on December 21st.

Note that the Sun rises due east and sets due west only on the first day of spring and the first day of autumn (about March 21st and September 22nd).
The Earth's axis of rotation is tilted 23.5 degrees to the plane of its orbit about the Sun.
When the northern hemisphere is tilted towards the Sun we have summer in the northern hemisphere and winter in the southern hemisphere.

When the northern hemisphere is tilted away from the Sun we have winter in the northern hemisphere and summer in the southern hemisphere.

This is the cause of the seasons.
Image of the Earth as it would appear from the Sun over the course of one year. Locations between +23 and -23 deg latitude see the Sun at the zenith twice a year at noontime.
The Cycle of the Seasons

1. You can use the celestial sphere to help you think about the seasons. The celestial sphere is the projection of Earth's spinning on the sky, and the equator is the projection of Earth's orbit on the sky. Because Earth is a sphere, the equator is a circle and the celestial equator is located at a latitude of 0°. As shown above, Earth's polar axis is tilted by 23.5°, as shown above. The sun moves between Earth's poles at the equinoxes, when the sun is directly overhead at the equator. The spring equinox is the moment of the spring equinox, and the autumnal equinox is the moment of the autumnal equinox. The winter solstice is the moment of the winter solstice, and the summer solstice is the moment of the summer solstice.

2. The two causes of the seasons are shown at right for a hemisphere. First, the sun's position on the celestial equator varies throughout the year, and the angle of the sun's rays varies throughout the year. Because of these variations, the amount of energy received by Earth from the sun varies throughout the year. The summer solstice is the moment of the summer solstice, and the winter solstice is the moment of the winter solstice.

3. At the summer solstice, the amount of energy received by Earth from the sun is at its maximum, and the amount of energy received by Earth from the sun is at its minimum. This is the reason why the summer solstice is the moment of the summer solstice, and the winter solstice is the moment of the winter solstice.

4. The winter solstice occurs on December 21 or 22, and the summer solstice occurs on June 20 or 21. The length of the day is shortest on the summer solstice, and the length of the day is longest on the winter solstice.

5. The spring equinox occurs on March 20 or 21, and the autumnal equinox occurs on September 22 or 23. The length of the day is equal on the spring equinox, and the length of the day is equal on the autumnal equinox.

6. The equinox occurs on March 20 or 21, and the vernal equinox occurs on March 20 or 21. The length of the day is equal on the vernal equinox, and the length of the day is equal on the autumnal equinox.

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The reason it is warmer in summer than in winter is that the sunlight is more concentrated on the ground when the Sun is higher in the sky. Also, in summer the Sun is above the horizon more hours per day.
The celestial poles (both north and south) are directly above Earth's poles.

Stars all appear to lie on the celestial sphere, but really lie at different distances.

The ecliptic is the Sun's apparent annual path around the celestial sphere.

The celestial equator is a projection of Earth's equator into space.
A model of the celestial sphere, showing the constellation borders, the celestial equator, and the apparent path of the Sun against the background of constellations.
The Sun is in the direction of the constellation Virgo in September, as viewed from the Earth. That is why people born in September are said to be born under the astrological sign of Virgo.
Just as we describe the location of a place on Earth by its latitude and longitude, we can specify the location of a star on the celestial sphere by its **right ascension** and **declination**.
About June 21st the Sun has a declination of +23.5 degrees. It is 23.5 degrees north of the celestial equator.

About March 21st and September 22nd the Sun is on the celestial equator and has a declination of 0.

About December 21st the Sun is 23.5 degrees south of the celestial equator. Its declination is -23.5 deg.
Because the Earth turns on its axis once a day, it appears that the stars move around the north and south celestial poles.

Sky at 40 deg N latitude
Stars close to the North Celestial Pole are always above the horizon. These are **circumpolar stars**.
In the northern hemisphere the circumpolar constellations move counterclockwise around the North Celestial Pole. In the southern hemisphere the circumpolar constellations move clockwise around the SCP.
The first regular observers of the sky (shall we say “astronomers”?) were the Chinese and the Babylonians. They divided up the sky into constellations or asterisms.

By modern agreement, the sky is divided into 88 constellations, some ancient, some relatively recent (18th century).

The twelve constellations of the zodiac (Capricornus, Aquarius, Pisces, Aries, Taurus, Gemini, Cancer, Leo, Virgo, Libra, Scorpius, and Sagittarius) are already familiar to you.
Hipparchus (ca. 140 BC) was perhaps the greatest astronomer of ancient times. He produced a catalogue of 1000 stars and classified them according to their apparent brightness.
The brightest stars were called stars of the first magnitude. Fainter stars were classified as being of second, third, fourth, or fifth magnitude. Now stars can be measured to +/- 0.01 magnitude. We have also expanded the scale to include negative values, and much larger positive values. Sirius, for example, has an apparent magnitude of -1.42. The faintest stars detectable with the Hubble Space Telescope are almost 30th magnitude.

If you measure the brightness of the stars with a photometer (a light measuring device), you will find that a 1st magnitude star gives 100 times as many photons as a 6th magnitude star. A 2nd magnitude star is 100 times more luminous than a 7th magnitude star...
Thus, the difference of two magnitudes is related to the ratio of the intensity of the light as follows:

\[ m_a - m_b = 2.5 \log \left( \frac{I_b}{I_a} \right) \]

If \( I_b = 100 \times I_a \), \( \log \left( \frac{I_b}{I_a} \right) = 2 \) and \( m_a - m_b = 5 \).

The astronomical magnitude scale is somewhat confusing, since fainter stars have larger magnitudes. But astronomers use magnitudes so much, we will have to get used to them. The best way to get familiar with magnitudes is to compare the stars on a star chart with real stars in the sky.
In order for a 5 magnitude difference to correspond to an intensity ratio of exactly 100, each magnitude actually corresponds to a factor of the fifth root of 100 (= 2.511886...) in light intensity.

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<thead>
<tr>
<th>Table 2-1</th>
<th>Magnitude and Intensity</th>
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α Orionis is also known as Betelgeuse.

β Orionis is also known as Rigel.
The brighter stars in a constellation are usually given Greek letters in order of decreasing brightness.

In Orion β is brighter than α, and κ is brighter than η. Fainter stars do not have Greek letters or names, but if they are located inside the constellation boundaries, they are part of the constellation.
Light curve of Betelgeuse (α Ori) from Oct. 21, 1979, to Nov. 11, 1996.

It is a slowly pulsating star that will eventually explode as a Type II supernova.
In order to describe the position of the Sun, Moon, stars, and planets in the sky, we need a coordinate system. The system we are most familiar with is the **horizon system**.

“Straight up” is called the **zenith**. The opposite point on the sky, which would be below your feet, is the **nadir**. These are the two poles of the horizon system. The horizon traces out a circle 90 degrees from these poles. For convenience we designate four **cardinal points** along the horizon: the north, east, south, and west points.

The number of degrees an object is above the horizon is called the “altitude” or **elevation angle**.
If we draw a line from the zenith through a celestial object and extend that line to the horizon, we obtain the **azimuth** angle of the object. By convention, the north point on the horizon has azimuth 0 degrees, the east point has azimuth 90 degrees, the south point has azimuth 180 degrees, and the west point has azimuth 270 degrees.

The problem with the horizon system is that the azimuth and elevation angle of a star changes continuously owing to the rotation of the Earth. If I said, “I saw a bright star at 30 degrees above the horizon in the east,” I would also have to specify the date and time of the observation, and my geographical coordinates, in order for someone to know for certain which star I was referring to.
Halfway between the NCP and the SCP is the celestial equator. It is a projection of the Earth's equator out to the celestial sphere.

The number of degrees that a celestial object is north or south of the celestial equator is called the declination (DEC). It is the analogue of latitude on the sky. The analogue of longitude is called right ascension (RA). While the RA and DEC of a star change slowly with time, these changes are very small fractions of a degree each year. Thus, we can make a star catalogue or star chart that is useful for observers at any location on the Earth. For example, the coordinates of Betelgeuse in the year 2000 were RA = 5 hours 55 minutes 10.3 seconds, DEC = +7 deg 24' 25".
It is also common to designate the right ascension by the Greek letter alpha (α) and the declination by the Greek letter delta (δ).

A third coordinate system uses the **ecliptic** as the fundamental great circle. This is the apparent path of the Sun in the sky against the background of stars. The Sun passes through all the constellations of the **zodiac** (plus Ophiuchus).
The word **planet** to the ancient Greeks meant “wandering star. We now know that they are other worlds. The orbital planes of the other planets are oriented quite similarly to that of the Earth. Thus, the planets are usually found within a few degrees of the ecliptic.
Hipparchus compared the coordinates of some stars with records made by the Babylonians and discovered that the ecliptic longitudes of the stars were increasing with time, about 1 degree per century. (The modern value is about 1 degree in 72 years.) This is called **precession** (not to be confused with the word “precision”). The Earth turning on its axis is like a spinning top.
While Polaris is close to the NCP now, it was not always the case. Due to the 26,000 year period of precession, many stars take their turns being the pole star.
Nutation ("nodding") of the axis of rotation is due to the tidal forces not being constant over time. 