Fourth-order algorithms for solving the imaginary-time Gross-Pitaevskii equation in a rotating anisotropic trap

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By implementing the exact density matrix for the rotating anisotropic harmonic trap, we derive a class of very fast and accurate fourth-order algorithms for evolving the Gross-Pitaevskii equation in imaginary time. Such fourth-order algorithms are possible only with the use of forward, positive time step factorization schemes. These fourth-order algorithms converge at time-step sizes an order-of-magnitude larger than conventional second-order algorithms. Our use of time-dependent factorization schemes provides a systematic way of devising algorithms for solving this type of nonlinear equations.

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I. INTRODUCTION

The dynamics of a fast rotating Bose-Einstein condensate (BEC) has been studied extensively in terms of the Gross-Pitaevskii (GP) equation [1,2]. By evolving the GP equation in imaginary time, it is easy to determine the ground state properties of the condensate, such as the formation of vortex-arrays and giant vortices [2,3]. It has been known for some time that the first-order pseudospectral, split-operator method [4] is a very fast way of solving the nonlinear Schrödinger equation. However, first or second-order split operator (SO) methods [5,6] and Crank-Nickolson (CN) algorithms with [7] or without [8] splitting ignore the time dependence of the nonlinear potential and converge linearly or quadratically only at very small time steps. Bandaruk and Shen [9] have applied higher order decomposition schemes with negative coefficients to solve the real time nonlinear Schrödinger equation. Due to the difficulty of estimating the nonlinear potential at intermediate time, they have not demonstrated that their higher order algorithms actually converge with accuracy beyond second order. In any case, their negative coefficient algorithms cannot be used for imaginary time evolution because negative time steps will result in an unbounded diffusion kernel [10–13].

In this work, we derive a class of very accurate fourth order factorization algorithm for solving the GP equation in imaginary time. These algorithms are made possible by the confluence of three key ideas: (1) The density matrix for a rotating anisotropic harmonic oscillator can be solved exactly. (2) The time dependence of the nonlinear potential can be systematically accounted for in factorization algorithms. (3) Forward, all positive time step algorithms [14–19] are now available for solving imaginary time evolution equations.

In the next section, we show how the density matrix of the harmonic oscillator can be exactly implemented as an algorithm. This obviates the need to expand in harmonic eigenstates [6,20]. In Sec. III, by exact diagonalization, we generalize the result to the case of a rotating anisotropic harmonic trap. In Sec. IV, we describe the time-dependent form of the fourth order forward algorithm for solving the GP equation. In Sec. V, we compare the convergence of various algorithms. We summarize our conclusions in Sec. VI.

II. EXACT ALGORITHM FOR THE HARMONIC OSCILLATOR

Consider the one-dimensional (1D) harmonic oscillator Hamiltonian operator given by

\[ H = T + V = \frac{1}{2} p^2 + \frac{1}{2} \omega^2 x^2. \]  

(2.1)

Its imaginary time propagator (or density matrix) can be exactly decomposed as

\[ e^{-\tau \hat{V}} \equiv e^{-\tau C_V} e^{-\tau C_T} e^{-\tau C_V}, \]

(2.2)

where \( C_V \) and \( C_T \) are functions of \( \tau \) to be determined. To show this, we simply compare the matrix elements on both sides. For the right hand side (rhs), we have (ignoring normalization factors)

\[ \langle x' | e^{-\tau C_V} e^{-\tau C_T} e^{-\tau C_V} | x \rangle \]

\[ = e^{-\tau \omega^2 x^2 / 2} e^{-(x'-x)^2 / (2 \tau \omega^2)} e^{-\tau \omega^2 x'^2 / 2}. \]  

(2.3)

For the left hand side (lhs) of (2.2), the exact density matrix element is known [21]

\[ \langle x' | e^{-\tau \hat{V}} | x \rangle \]

\[ = \exp \left( -\frac{\omega}{2 \sinh(\omega \tau)} \left( (x'^2 + x^2) \cosh(\omega \tau) - 2 x' x \right) \right) \]

\[ = \exp \left( -\frac{\omega}{2 \sinh(\omega \tau)} \left( (x'^2 + x^2) \cosh(\omega \tau) - 1 \right) + (x' - x)^2 \right), \]

(2.4)

where we have expressed \(-2 x' x = (x' - x)^2 - x'^2 - x^2\). Compar-
ing (2.3) to (2.4) allows us to identify the coefficient functions as

\[
C_v = \frac{\cosh(\omega \tau) - 1}{\omega \tau \sinh(\omega \tau)} \quad \text{and} \quad C_T = \frac{\sinh(\omega \tau)}{\omega \tau}.
\]

In the limit of \(\tau \to 0\), we have

\[
C_v = \frac{1}{2} - \frac{1}{24} \omega^2 \tau^2 + \frac{1}{576} \omega^4 \tau^4 + \cdots,
\]

\[
C_T = 1 + \frac{1}{6} \omega^2 \tau^2 + \frac{1}{192} \omega^4 \tau^4 + \cdots.
\]

If we keep only the first term, we have a second-order algorithm. Keeping the first two terms gives a fourth-order algorithm, keeping the first three terms gives a sixth-order algorithm, etc.

The exact factorization (2.2) is possible for the harmonic oscillator because its Hamiltonian is quadratic and higher order commutators are either zero or simply proportional to \(T\) or \(V\). The harmonic oscillator is characterized by two key commutators,

\[
[V,[T,V]] = 2\omega^2 V,
\]

\[
[T,[V,T]] = 2\omega^2 T.
\]

Because of these two equalities, all higher order commutators can be subsumed back to the original operators \(T\) and \(V\). To see how this exact decomposition comes about, let us begin with the simple second-order decomposition

\[
e^{-\tau(T+V)} = \exp\left[-\tau(T+V) + \frac{1}{2\tau} \tau^3 [V,[T,V]] - 2[T,[V,T]] + O(\tau^5]\right].
\]

Since the error terms are proportional to the original operators, they can be symmetrically moved back to the lhs to yield,

\[
e^{-\tau(T+V)} = e^{-\tau(T+V) + O(\tau^5)}.
\]

The decomposition of the lhs is then correct to the fourth order. The coefficients agree with the expansion (2.6) and (2.7). This example makes it clear that the exact expansion only depends on the abstract commutator relations (2.8) and (2.9), and is independent of the specific representation of the 1D harmonic oscillator. Also, if we exchange the operators \(T \leftrightarrow V\), the coefficients are unchanged. Thus we can also factorize exactly via

\[
e^{-\tau(T+V)} = e^{-\tau V T} e^{-\tau V} e^{-\tau T}.
\]

For real time propagation, we only need to set \(\tau = it\) to get the corresponding coefficients,

\[
C_v = \frac{1 - \cos(\omega t)}{\omega t \sin(\omega t)} \quad \text{and} \quad C_T = \frac{\sin(\omega t)}{\omega t}.
\]

For either real or imaginary time evolution, one iterates the discretized wave function forward in time via

\[
\psi(\tau + \Delta \tau) = e^{-\Delta \tau T} \psi(\tau).
\]

If the exact density matrix (2.4) were used directly in coordinate space, that would incur a slow, \(N \times N\) matrix multiplication of the Gaussian kernel \(e^{-\tau \sigma (t')^2}\). The advantage of the factorized form is that this matrix multiplication can be avoided by going to \(k\) space via fast Fourier transform (FFT) and multiplying the \(k\)-space wave function point-by-point by \(e^{-\tau \sigma (t')^2}\). This is then an order \(N \ln_2 N\) operation, much faster than the \(N \times N\) coordinate space matrix multiplication.

### III. Exact Algorithm for a Rotating Anisotropic Harmonic Trap

Consider now the case of an rotating anisotropic harmonic potential with Hamiltonian

\[
H = \frac{1}{2}(p_x^2 + p_y^2) + \frac{1}{2} \omega_0^2 x^2 + \frac{1}{2} \omega_y^2 y^2 - \tilde{\Omega}(x p_y - y p_x).
\]

This is a well-studied problem in nuclear physics [22]. Its diagonalization is greatly simplified [23] if we characterize the anisotropy via the deformation parameter \(\delta\),

\[
\omega_0^2 = (1 + \delta) \omega_0^2, \quad \omega_y^2 = (1 - \delta) \omega_0^2,
\]

measure lengths in units of the oscillator length \(l = 1 / \sqrt{\omega_0}\), and express \(H\) and \(\tilde{\Omega}\) in units of \(\omega_0\). The resulting dimensionless Hamiltonian is then

\[
H = \frac{1}{2}(p_x^2 + p_y^2) + \frac{1}{2} (1 + \delta) x^2 + \frac{1}{2} (1 - \delta) y^2 - \Omega(x p_y - y p_x),
\]

where \(\Omega = \tilde{\Omega} / \omega_0\). To diagonalize this Hamiltonian, we introduce two new sets of canonical variables

\[
Q_1 = \alpha_1 (c x - s y), \quad P_1 = \frac{1}{\alpha_1} (c p_x + s y),
\]

\[
Q_2 = \alpha_2 (c y - s x), \quad P_2 = \frac{1}{\alpha_2} (c p_y + s x),
\]

where \(\alpha_i\) are normalization constants, and \(c = \cos(\phi), s = \sin(\phi)\). One can check that the canonical commutator relations are indeed satisfied

\[
\{Q_i, P_j\} = i \delta_{ij}.
\]

In terms of \(\{Q_i, P_j\}\), because of the way we have parameterized the anisotropy and expressed everything in terms of \(\omega_0\), the coefficients of both \(P_i Q_j\) and \(P_j Q_i\) can be made to vanish with a single condition

\[
\tan(2 \phi) = \frac{2 \Omega}{\delta}.
\]

Using \(\alpha_i\) to normalize the \(P_i^2\) terms with unit coefficient, the resulting Hamiltonian can be written as
where
\[
\alpha_1^2 = 1 - \frac{6}{\delta} \frac{1}{\sqrt{\delta^2 + 4(\delta^2)}},
\]
(3.8)
and
\[
\alpha_2^2 = 1 + \frac{6}{\delta} \frac{1}{\sqrt{\delta^2 + 4(\delta^2)}},
\]
(3.9)
with
\[
\Omega_1^2 = 1 + \Omega^2 - \sqrt{\delta^2 + 4(\delta^2)}.
\]
\[
\Omega_2^2 = 1 + \Omega^2 + \sqrt{\delta^2 + 4(\delta^2)}.
\]
(3.10)
Also, from (3.7), we have
\[
2\psi = 1 - \frac{\delta}{\sqrt{\delta^2 + 4(\delta^2)}},
\]
(3.11)
\[
2\psi^2 = 1 + \frac{\delta}{\sqrt{\delta^2 + 4(\delta^2)}}.
\]
At \(\Omega = 0\), the phase angle \(\phi = 0\). As \(\Omega\) increases, the phase angle approaches 45° asymptotically. Thus \(x\) and \(c\) in (3.11) are both positive. However, as \(\Omega\) increases, \(\Omega_2^2\) crosses zero and becomes negative at \(\Omega = \sqrt{1-6}\). At this critical rotation rate, the Coriolis force overcomes the weaker harmonic potential in the \(\psi\) direction and the anisotropic harmonic oscillator is unstable. \(\Omega_2^2\) emerges positive again when \(\alpha_2^2\) crosses zero and turn negative at \(\Omega^2 = 1 + \delta\). Thus \(\Omega_2^2\) is negative over the interval \(1 - \delta \leq \Omega^2 \leq 1 + \delta\). This is an instability of the rotating harmonic oscillator, not necessary that all of the Gross-Pitaevskii equation. We will come back to this point in Sec. VI. Note also that for \(\delta = 0\), the algorithm is stable up to \(\Omega = 1\).

Equation (3.8) consists of two independent harmonic oscillators with different frequency. The two exact algorithms must be applied in sequence. However, since \(T_1\) and \(V_2\) only depend on \(p_x\) and \(y\), they should be placed next to each other so that both can be evaluated in the same mixed representation described below. Similarly, \(T_2\) and \(V_1\) only depend on \(x\) and \(p_y\). We therefore use the following factorization for each algorithm:
\[
e^{-\pi T_1 \psi V_1} = e^{-\pi C_T(1)T_1} e^{-\pi \gamma V_1} e^{-\pi C_T(1)T_1},
\]
(3.12)
\[
e^{-\pi T_2 \psi V_2} = e^{-\pi C_T(2)V_2} e^{-\pi \gamma V_2} e^{-\pi C_T(2)V_2},
\]
(3.13)
and interlaced them as follows:
\[
e^{-\pi T_1 \psi V_1 + \psi V_2} = e^{-\pi C_T(2)V_2} e^{-\pi \gamma V_2} e^{-\pi C_T(2)V_2}. \]
(3.14)
Here we use the shorthand notations \(C_T(1) = C_T(\Omega_1)\), \(C_T(2) = C_T(\Omega_2)\), etc. To implement (3.14), let
\[
\psi(x,y) = \frac{1}{\sqrt{2\pi}} \int dp_x \phi(p_x, y) e^{ip_x x},
\]
(3.15)
\[
\phi(p_x, y) = \frac{1}{\sqrt{2\pi}} \int dx \psi(x, y) e^{-ip_x x},
\]
(3.16)
and
\[
\phi(p_x, y) = \frac{1}{\sqrt{2\pi}} \int dy \psi(x, y) e^{-ip_y y}.
\]
(3.17)
The operators \(T_1\) and \(V_2\) are diagonal in the representation \(\phi(p_x, y)\) and \(T_2\) and \(V_1\) are diagonal in the representation \(\psi(x,y)\). In practice, \(\psi(x, y)\) is discretized as an \(N\times N\) complex array and its Fourier transform is computed using the discretized FFT. Thus the exact algorithm consists of four steps:

1. Compute the forward \(N\times 1\) transform \(\phi(p_x, y)\) from \(\psi(x, y)\) and multiply \(\phi(p_x, y)\) grid point by grid point by \(e^{-\pi C_T(1)T_1 - \pi C_T(2)V_2}\), where \(T_1\) and \(V_2\) are now understood to be functions of \(p_x\) and \(y\).

2. Compute the two-dimensional (2D) transform \(\phi(x, p_y)\) from the updated \(\phi(p_x, y)\) and multiply \(\psi(x, p_y)\) by \(e^{-\pi C_T(1)T_1 - \pi C_T(2)V_2}\), where \(T_1\) and \(T_2\) are now functions of \(x\) and \(p_y\).

3. Compute the inverse 2D transform from the updated \(\phi(x, p_y)\) back to \(\psi(p_x, y)\) and multiply \(\psi(p_x, y)\) by \(e^{-\pi C_T(1)T_1 - \pi C_T(2)V_2}\).

4. Compute the backward \(N\times 1\) transform from the updated \(\phi(p_x, y)\) back to \(\psi(x, y)\).

Thus the algorithm can be implemented with only three 2D FFTs. (One 2D transform=2N 1D transforms.) This is only one 2D FFT more than solving the nonrotational case.

IV. Solving the Gross-Pitaevskii Equation

Denoting now the entire rotating trap Hamiltonian (3.8) as the operator
\[
\hat{T} = T_1 + V_1 + T_2 + V_2.
\]
the corresponding 2D Gross-Pitaevskii equation is
\[
(T + g |\psi|^2) \psi(x, y) = \mu \psi(x, y).
\]
The condensate ground state can be projected out by imaginary time evolution
\[
\psi_0 \propto \lim_{\tau \rightarrow -\infty} \psi(\tau) = \lim_{\tau \rightarrow -\infty} e^{-\pi (T + V(\tau)) + \pi \mu} \phi(0).
\]
The chemical potential \(\mu\) is determined by preserving the wave function’s normalization to unity. This will be taken for granted and this term will be ignore in the following discussion. Since \(\psi(\tau)\) is time dependent, we have explicitly indicated that the Gross-Pitaevskii potential

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The calculation uses 642 grid points over a 142 harmonic coupling strength and not close to any particular limit. This choice corresponds to sizable anisotropy, rotation, catastrophically. Thus if one approximates

$$V(\tau) = g |\psi(\tau)|^2,$$  
(4.4)

is also time dependent.

In general, to solve (4.3) by factorization algorithms, one must apply rules of time-dependent factorization [16,24]: the time-dependent potential must be evaluated at an intermediate time equal to the sum of time steps of all the T operators to its right. For example, the first-order algorithm 1A is

$$\psi(\Delta \tau) = e^{-\Delta \tau T} e^{-\Delta \tau V(0)} \psi(0)$$  
(4.5)

and the first-order algorithm 1B is

$$\psi(\Delta \tau) = e^{-\Delta \tau V(\Delta \tau)} e^{-\Delta \tau T} \psi(0).$$  
(4.6)

While algorithm 1A is straightforward, 1B requires that the potential be determined from the wave function to be computed. This self-consistency condition can be solved by iterative methods described below.

In contrast to real time propagation, the wave function in imaginary time converges quickly to an approximate ground state depending on $\Delta \tau$ and produces a $\psi(\Delta \tau)$ that differs from $\psi(0)$ only by a normalization constant. Thus after some initial iterations, the normalized $g(\psi(\Delta \tau))^2$ is independent of $\tau$ and can be replaced by $g(\psi(0))^2$. This replacement can be justified only at small $\Delta \tau$ when the approximate wave function is close to the exact ground state and is unchanging in time. (For real time propagation, the wave function is always changing with time, and one cannot justify this replacement even at small $\Delta \tau$.) At larger $\Delta \tau$, the approximate ground state may not be a discrete bound state and the algorithm may fail catastrophically. Thus if one approximates $g(\psi(\Delta \tau))^2$ by $g(\psi(0))^2$ in (4.6), then the algorithm is still first order, but only at very small $\Delta \tau$. We shall refer to this version of the algorithm as 1B0.

We define the second-order algorithm 2A as

$$\psi(\Delta \tau) = e^{-2(1/2)\Delta \tau V(\Delta \tau)} e^{-\Delta \tau T} e^{-2(1/2)\Delta \tau V(0)} \psi(0)$$  
(4.7)

and algorithm 2B as

$$\psi(\Delta \tau) = e^{-2(1/2)\Delta \tau V(\Delta \tau)} e^{-\Delta \tau T} \psi(0).$$  
(4.8)

Similarly, one can replace $g(\psi(\Delta \tau))^2$ by $g(\psi(0))^2$ in algorithm 2A without affecting its quadratic convergence at very small $\Delta \tau$. We shall refer to this version of the algorithm as 2A0. Algorithm 2B requires two executions of the exact algorithm (3.14) for similar convergence, which is less efficient. We therefore did not implement algorithm 2B.

Figure 1 shows the convergence of algorithm 1A and 1B0 for the chemical potential $\mu$. Both are very linear at small $\Delta \tau$. The calculation is done for $\delta=0.5$, $\Omega=0.5$, and $g=50$. This choice corresponds to sizable anisotropy, rotation, coupling strength and not close to any particular limit. The calculation uses 642 grid points over a 142 harmonic length square centered on the origin. Changing the grid size to 1282 only changes the stable results in the fifth or sixth decimal place. The ground state wave function is nearly converged by $\tau=2$. The chemical potential shown is calculated at $\tau=10$. Note that the linear convergence line for 1B0 fails abruptly at $\Delta \tau=0.15$. Since algorithm 2A0 is just running algorithm 1A first followed by 1B0 at half the time-step size, the convergence failure of 1B0 accounts for the failure of algorithm 2A0 near $\Delta \tau=0.3$. Both algorithm 1A and 1B0 require an exceedingly small $\Delta \tau(<0.001)$ to produce an accurate value of $\mu$. Even algorithm 2A0 requires $\Delta \tau$ to be $<0.05$.

V. SELF-CONSISTENT ITERATIONS

To see the full effect of time-dependent factorization, we must implement 1B and 2A in the form (4.6) and (4.7) with self-consistent iterations to determined the GP potential. The required consistency equation is of the form

$$\psi = \frac{1}{\sqrt{Z}} e^{-b(1/2)|\phi|^2} \phi,$$  
(5.1)

where $b=c\Delta \tau g$ for some coefficient $c$, $\phi$ is the unnormalized intermediate wave function prior to the evaluation of the potential term, and $Z$ is the constant that normalizes $\psi$. This is needed because we are solving for the GP potential, which requires a normalized wave function. Since only the square of the modulus is needed, we solve (5.1) as

$$x = \frac{1}{Z} e^{-b|\phi|^2} \phi$$  
(5.2)

where $x=|\psi(r)|^2$, $a=|\phi(r)|^2$ and

$$Z = \int e^{-b|\phi|^2} d\mathbf{r} = \sum_i e^{-b|\phi_i|^2} (\Delta x)^2.$$  
(5.3)

It is helpful to view these equations in the discrete forms in which they are actually solved. When necessary, we will denote array elements explicitly as $x_i=|\phi_i(r)|^2$, etc.
A naive way of solving (5.2) is just to iterate
\[ x_{n+1} = F(x_n) = \frac{1}{Z(x_n)}e^{-b x_n a}. \] (5.4)
Starting with \( x_0 = 0 \), the first iteration would produce the normalized \( \bar{a} = a/Z \), which is a reasonable starting guess. Denoting \( x_n = x^* + e_n \), where \( x^* \) is the exact solution, then
\[
\begin{align*}
e_{n+1} &= F'(x^*)e_n + O(e_n^2) \\
&= -b x^* (1 - O(1/N))e_n + \cdots \\
&= -(c \Delta \tau g x^*)e_n. \tag{5.5}
\end{align*}
\]
The \( O(1/N) \) term neglected above is from differentiation 1/Z with respect to \( x_i = e^{-b(x_n)q_i/\Sigma_{j=1}^N e^{-b(x_n)q_j}} \), where \( N \) is the total number of array elements. It is therefore an excellent approximation to regard \( Z \) as a constant when differentiating with respect to individual array elements. The error propagation (5.5) explains why this naive iteration is undesirable; it will diverge abruptly at large \( \Delta \tau \) such that \( |c \Delta \tau g x^*| > 1 \).

The self-consistency equation can be solved by better methods, such as Newton’s or Halley’s iterations. However, at large \( \Delta \tau \), the normalized \( \bar{a} \) is not a good enough starting point. Fortunately, (5.2) has the exact solution
\[
x = \frac{1}{b} W\left(\frac{ba}{Z}\right) \tag{5.6}
\]
where \( W(y) \) is the Lambert \( W \) function [25]
\[
W e^W = y \tag{5.7}
\]
with series expansion
\[
W(y) = y - y^2 + \frac{3}{2} y^3 - \frac{8}{3} y^4 + O(y^5). \tag{5.8}
\]
The series expansion (5.8) is not useful for our purpose since its radius of convergence is only \( 1/e \). A better choice is the following uniform approximation by Winitzki [26]:
\[
W(y) \approx \ln(1+y) \left\{ 1 - \frac{\ln[1 + \ln(1+y)]}{2 + \ln(1+x)} \right\}, \tag{5.9}
\]
which has a maximum relative error of 2% (at \( x = 2 \)) in \([0, \infty]\).

The normalization constant \( Z \) can in principle be determined from
\[
b = \int W\left(\frac{ba}{Z}\right) d\tau, \tag{5.10}
\]
but this integral equation is time consuming to solve when \( W(y) \) itself has to be computed numerically. A more workable scheme is to compute \( Z \) via (5.3) from \( x_n \), and compute \( x_{n+1} \) via (5.6). The use of (5.6) will guarantee convergence at all \( b \) and \( \Delta \tau \), provided that one can compute \( W(y) \) in a simple way. Thus we use the normalized \( \bar{a} \) as the starting value, compute \( Z \) via (5.3), solve for \( x \) via (5.6) and normalize it to obtain an approximate GP potential at all \( \Delta \tau \). We find that this one iteration is sufficient to remove all instability in algorithms 1A and 1B; further Newton-Raphson iterations produce marginal improvements not worth the additional effort.

Figure 1 shows the convergence of algorithms 1B and 2A when the self-consistency condition is approximately satisfied by our one \( W \)-function iteration. The results are denoted as 1BW and 2AW. The instability in 1B0 and 2A0 no longer appears. The convergence of both are linear and quadratic out to large values of \( \Delta \tau \).

For first- and second-order algorithms, self-consistent iterations are not needed because \( \Delta \tau \) has to be very small in order for these algorithms to produce results close to the exact one. If \( \Delta \tau \) is small, then one may as well use 1B0 and 2A0 without wasting time on self-consistent iterations. Self-consistency is a concern only when one is interested in enlarging the step-size convergence of higher order algorithms.

VI. FOURTH-ORDER ALGORITHMS

It is well known from studies of symplectic integrators that factorizations of the form (4.5)–(4.8) can be generalized to higher order in the form [27–34]
\[
e^{-\Delta (\tau + \nu)} = \prod_i e^{-a_i \Delta \tau} e^{-b_i \Delta \nu}, \tag{6.1}
\]
with coefficients \( \{a_i, b_i\} \) determined by the required order of accuracy. However, as first proved by Sheng [35] and made explicit by Suzuki [36], beyond second order, any factorization of the form (6.1) must contain some negative coefficients in the set \( \{a_i, b_i\} \). Goldman and Kaper [37] later proved that any factorization of the form (6.1) must contain at least one negative coefficient for both operators. Since a negative time step for the kinetic energy operator will result in an unbound and unnormalizable wave function, no such factorization scheme can be used to evolve the imaginary time Schrödinger equation, including the Gross-Pitaevskii equation. To go beyond second order, one must use forward factorization schemes with only positive factorization coefficients [14–16]. These forward algorithms are currently the only fourth-order algorithms possible for solving time-irreversible equations with a diffusion kernel [10,11] and have been applied successfully in solving the imaginary time Schrödinger equation [12,13]. Omelyan, Mryglod, and Folk [18,19] have compiled an extensive list of fourth and higher order symplectic algorithms. However, their sixth- and eighth-order algorithms contain negative time steps and are not forward algorithms. Recently, one of us has proved [38] that while sixth-order forward algorithms are possible, they require an additional commutator currently not implementable. Thus forward algorithms are very unique. Here, we will show that they also yield highly accurate fourth-order algorithms for solving the Gross-Pitaevskii equation.

The problem we seek to solve is the ground state of
\[
H = H_s + H_v + V(x, y, \tau) \tag{6.2}
\]
where \( V(x, y, \tau) \) is the GP potential (4.4) and
\[
H_s = \frac{1}{2} \rho_x^2 + \frac{1}{2} (1 + \delta) x^2 - \Omega x p_y, \tag{6.3}
\]
The Hamiltonian fundamentally has three operators, which are diagonal in \((x, y), (x, p_x), \) and \((p_x, y)\). If the external trapping potential \(V_{\text{ext}}(x, y)\) is more general and nonharmonic, we can still write
\[
H = H_x + H_y + V(x, y, \tau)
\]
but now with
\[
V(x, y, \tau) = V_{\text{ext}}(x, y) - \frac{1}{2}(1 + \delta)\Delta x^2 - \frac{1}{2}(1 - \delta)\Delta y^2 + g|\phi(x, y, \tau)|^2.
\]

The parameter \(\delta\) is then a free parameter associated with algorithm, which we can choose to match the asymmetry of \(V_{\text{ext}}\), or just set to zero. The crucial points is that, for a rotating trap, harmonic or not, the Hamiltonian has three operators diagonal in three separate spaces. By computing the density matrix of
\[
T = H_x + H_y
\]
exactly via algorithm (3.14), we have reduced the Hamiltonian to a two-operator problem. This is a tremendous simplification. This simplification is not restricted to harmonic traps, but holds equally for an arbitrary external potential. The key point is that the rotating part of the Hamiltonian can be diagonalized regardless of the choice of the confining potential. When we diagonalize \(H_x + H_y\), we generate an inverted harmonic potential in (6.6), which must be compensated by the external potential or the GP potential. In the following we will present results only for (6.2), but our algorithm works in the general case of (6.6). We will come back to this point when we discuss overcritical rotation in the next section.

Because implementing \(T\) is computational demanding, we must choose a fourth-order algorithm with a minimal number of \(T\) operators. Thus among the many forward algorithms discovered so far [15–19], we choose to implement only the simplest algorithm, 4A,
\[
\phi(\Delta \tau) = e^{-(1/6)\Delta \tau^2(\Delta \tau)} e^{-(1/2)\Delta \tau T} e^{-(2/3)\Delta \tau^2(\Delta \tau)} e^{-(1/2)\Delta \tau^2 V(0)} \phi(0),
\]
with \(\tilde{V}\) given by
\[
\tilde{V} = V + \frac{\Delta \tau^2}{48} [V, [T, V]].
\]

Despite the seeming complexity of \(T\) as defined by the Hamiltonian (3.8), we have remarkably
\[
[V, [T, V]] = \left( \frac{\partial V}{\partial x} \right)^2 + \left( \frac{\partial V}{\partial y} \right)^2.
\]

Thus, the midpoint effective potential is
\[
H_y = \frac{1}{2} \Delta \tau^2 + \frac{1}{2} (1 - \delta)\Delta y^2 + \Omega y p_x.
\]
All the fitted lines are of the monomial form
\[ E = E_0 + C \Delta \tau^p. \] (6.15)

Algorithms 1A and 1B0 yielded near-identical quadratic convergence. Algorithm 2A0 can be fitted with a fourth order monomial as shown. The fit is not perfect because because our W function is only an approximation. Algorithm 4A00 and 4A0W failed too abruptly to show a smooth trend, but 4AWW can indeed be fitted with an eighth-order monomial.

The computational effort required by each algorithm is essentially that of evaluating the exact algorithm (3.14), which uses three 2D FFT. Since 1A, 1B0, and 2A0 all use the exact algorithm once, the second-order algorithm 2A0 is clearly superior. Algorithms 4A00 requires two evaluations of the exact algorithm plus the gradient potential. The gradient potential, if done by FFT, requires two 2D FFT. Thus algorithm 4A00 requires eight 2D FFT, which is \(8/3=3\) times the effort of algorithm 2A0. Since algorithm 4A00 converges much better than 2A0 at time steps more than three times as large, the class of 4A00 and 4A0W algorithm is clearly more efficient. This efficiency is especially evident if higher accuracy is required. The fully implemented algorithms 4AWW use the second-order algorithm to evaluate midpoint wave function and is therefore \(\approx 2\) times the effort of 2A0. Looking at Fig. 2, algorithms 4AWW clearly converge better than 2A0 (2A0) even at time steps four times as large. Note that the first- and second-order algorithms are basically similar, whereas all fourth-order algorithm are qualitatively distinct. The second-order algorithm is not an order-of-magnitude better than a first-order algorithm, whereas all fourth-order algorithms are an order-of-magnitudes better than the second-order algorithm.

This advantage of fourth-order algorithms is cumulative. For example, one can quickly evolve into the ground state by use of large time steps. As stated earlier, the GP ground state can be obtained at \(\tau\approx2\). Using algorithm 4A0W at \(\Delta \tau=0.5\), one can get there in four iterations. Algorithm 2A0 would have taken 80 iterations at \(\Delta \tau=0.02\).

To see how these comparisons work out in practice, we give below some timing information. Since running time is code and machine dependent, this should be viewed as merely illustrative. The algorithms were programmed in Fortran 90 and ran on a 1.2 GHz Pentium machine with 0.5 GB of RAM. The IMSL 2D FFT is used. For the case of a 64 × 64 point mesh, each run of algorithms 1A, 1B0, 2A0, 1BW, and 2AW required 0.0130, 0.0135, 0.0138, 0.0173, and 0.0203 sec, respectively. (Timing is obtained by averaging over 500 run of each algorithm.) Algorithms 1A, 1B0, and 2A0 are indeed comparable, but each self-consistency iteration increases the time by \(0.006\) sec, which is an increase of 40% for 2A0. The time for algorithm 4A00, 4A0W, and 4AWW are, respectively, 0.0394, 0.0459, and 0.0630 sec. The time for algorithm 4A00 is approximately three times that of 2A0 and the self-consistency iteration now accounts for only an 16% increase. Algorithm 4AWW is slightly more than four times that of 2A0. In the case of a 128 × 128 point mesh, the timing for 2A0, 4A00, 4A0W, and 4AWW are, respectively, 0.0568, 0.1669, 0.1916, and 0.2617 sec. The ratio of timing remained similar. For three and four times the effort, these fourth-order algorithms can converge at time steps at least 20 times as large as second-order algorithms. Moreover these are \(O(N)\) algorithms which scale with the mesh size. When doing a three-dimensional (3D) \(N=(256)^3\) calculation, the required time would be thousands times longer and the gain in using fourth-order algorithms would be even more measurable.

By solving the density matrix of the rotation harmonic oscillator (6.7) exactly, we have effected a tremendous simplification which has allowed us to derive very compact fourth-order algorithms with excellent large time-step convergence. They are no more difficult to implement than second-order algorithms. If we do not have the exact density matrix, then we would have to approximate each occurrence of \(e^{-i(1/2)\Delta \tau^2}\) in (6.8) to fourth order, resulting in a much more complex algorithm.

However, by solving the rotating harmonic oscillator exactly, the current algorithms also inherited its limitations. As alluded to in Sec. III, the rotating harmonic trap becomes unstable at \(\Omega_\ast=\sqrt{1-\delta}\). Thus if we are to use the exact algo-
rithm (3.14), we must require $\Omega < \Omega_c$. However, it is known [39,40] that for $\delta < 1/5$ and $g$ sufficiently large, the full GP equation can support “overcritical” rotation in the interval $\sqrt{1-\delta} < \Omega < \sqrt{1+\delta}$. For overcritical rotation, one must not prematurely impose the limitation of the rotating harmonic oscillator on the algorithm.

In light of our previous discussion, we now consider the case of

$$V_{ex}(x,y) = \frac{1}{2}(1+\Delta)x^2 + \frac{1}{2}(1-\Delta)y^2,$$

and group the Hamiltonian as follows:

$$H = H_i + H_v + V(x,y,\tau),$$

(6.17)

where $H_i$ and $H_v$ are defined as before in (6.3) and (6.4), and

$$V(x,y,\tau) = \frac{1}{2}(\Delta-\delta)x^2 - \frac{1}{2}(\Delta-\delta)y^2 + g|\psi(x,y,\tau)|^2.$$

(6.18)

We thus divorce the deformation parameter $\delta$ associated with the algorithm, from the physical deformation parameter $\Delta$ associated with the trapping potential. If we choose $\delta=0$, the algorithm is stable up to $\Omega=1$, above the physical critical value of $\Omega_c = \sqrt{1-\Delta}$. Moreover, since in the Thomas-Fermi approximation the density profile follows the shape of the potential, (6.18) indeed suggests that the inverted harmonic potential $-\frac{1}{2}\Delta y^2$ can be compensated by the GP potential at sufficiently large $g$, making overcritical rotation possible.

### VII. CONCLUDING SUMMARY

In this work we have derived a number of fourth-order algorithms and demonstrate their fourth-order convergence in solving the GP equation in a rotating, anisotropic harmonic trap. These fourth-order algorithms, based on forward factorization schemes, are the only class of factorization algorithms possible for solving evolution equations with a diffusion kernel. Our use of the time-dependent factorization rule provided a systematic way of solving the nonlinear GP equations and can be generalized to solve similar nonlinear equation such as the Hartree-Fock and the Kohn-Sham equation [41,42]. These fourth-order algorithms are particularly efficient in solving for the ground state by use of large time steps. In contrast to other algorithms, generalizing these algorithms to 3D is very transparent, one simply replaces 2D FFT everywhere by 3D FFT. Our use of the exact algorithm, which diagonalizes the rotating component of the Hamiltonian, is general and can applied to any external trapping potential. This exact algorithm also provided insight for understanding overcritical rotation. Physical results obtained by applying these algorithms will be presented elsewhere.

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